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Planetary Low Tide May Force Regular Sunspot Sync Ups

or more than 1,000 years, the number of sunspots hit a minimum within a few years of a major planetary alignment. A recent study showed that tides created by this alignment every 11 years are strong enough to tug on material near the Sun's surface and synchronize localized changes in its magnetic field.

"We noticed from historical data that there is an astonishing degree of regularity" in the sunspot cycle, Frank Stefani, lead author of the study, told *Eos.* Stefani is a fluid dynamics research fellow at Helmholtz-Zentrum Dresden-Rossendorf in Dresden, Germany. "We definitely have a clocked process," he said. "But then the question was, What is the clock?"

The study expands upon the commonly accepted model for the solar dynamo and supports a long-held theory that planetary configurations are responsible for the sunspot cycle and magnetic solar cycle.

Other planetary systems might have tidally dominant planets that resonate with their suns, but it's not likely that we'll be able to prove it.

Wound, Twisted, and Unstable

As a giant spinning ball of plasma, the Sun's magnetic field is extremely complicated. Its magnetic field lines start as parallel lines running from the north pole to the south. But because the Sun rotates faster at its equator than at its poles, those pole-to-pole magnetic field lines slowly wind and wrap around the Sun, stretching like taffy from the middle of the line to become horizontal.

In addition to the rotational motion of solar plasma, convection moves material from the equator to the poles and back again. This twists the field lines around each other into loops and spirals.

The winding and twisting of the Sun's magnetic field lines are described by the alphaomega dynamo model. In that model, alpha represents the twisting, and omega represents the wrapping. Tangled field lines can create instabilities in the local magnetic field and cause sunspots, flares, or mass ejections.

This model is the commonly accepted explanation for the behavior of the Sun's magnetic field, but it's not perfect, Stefani explained. It predicts that the instabilities' twistedness will oscillate randomly every few years. But the model can't explain why the number of sunspots waxes and wanes on a roughly 11-year cycle or why the Sun's magnetic field flips polarity every 22 years.

Low Tide, Low Activity

Another solar system phenomenon happens every 11 years: Venus, Earth, and Jupiter align in their orbits. These three planets have the strongest tidal effect on the Sun, the first two because of their proximity to the Sun and the third because of its mass. Past observational studies have shown that minima in the sunspot cycle have occurred within a few years of this alignment for the past 1,000 or so years.

"If you look at the trend, it has an amazing parallelism," Stefani said.

The researchers wanted to test whether the planetary alignment could influence the Sun's alpha effect and force an interplanetary low tide at regular intervals. They started with a standard alpha-omega dynamo model and added a small tidal tug to the alpha effect every 11 years to simulate the alignment.

"Our dynamo model is not a completely new one," Stefani explained. "We're really building on the old-fashioned, or conventional, alpha-omega dynamo."

The simulation showed that even a weak tidal tug of 1 meter per second every 11 years forced unstable magnetic twists to pulse with that same period. The simulated dynamo's



An ultraviolet image of the Sun from 2016 is overlaid with a map of its magnetic field lines. Credit: NASA/SDO/AIA/LMSAL

polarity oscillated with a 22-year period, just like the real solar dynamo.

"With a little bit of this periodic alpha," Stefani said, "we can indeed synchronize the dynamo period to 22 years [with] planetary forcing."

Because those magnetic instabilities are connected with solar activity, the researchers argue, this synchronization could also suppress (or generate) sunspots across the Sun at roughly the same time—in other words, the sunspot cycle. The team published these results in *Solar Physics* in late May (see bit.ly/ dynamo-model).

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A Counterintuitive Result?

"This is an intriguing paper," said Steve Tobias, a solar dynamo researcher at the University of Leeds in the United Kingdom who was not involved with the research. Tobias argued that the combined planetary tides are too weak to directly set the length of the solar cycle—plasma dynamics deep within the Sun are the more likely cause, he told *Eos*.

Nevertheless, this study "seems to show that even a tiny amount of forcing from tidal processes can resonantly synchronize the cycle," Tobias said. "This counterintuitive result should be explored further by investigating the behavior of proxies for solar activity, such as the production rates of isotopes of beryllium deposited in ice cores."

It's possible that other planetary systems might have tidally dominant planets that resonate with their suns like ours do, Stefani said, but it's not likely that we'll be able to prove it.

For most stars, "we have observations going back about 40 years," he said. "And people are happy if they can identify two or three or four periods. Only for our Sun do we have all the



A simplified schematic of a single magnetic field line as it wraps around the Sun (the omega effect) and then twists upon itself (the alpha effect). The arrows indicate the direction that solar material moves as it drags the field line with it. Credit: NASA/MSFC

historical observations. We have beryllium data. We can go back for thousands of years."

"Our Sun is quite an ordinary star, but it is quite special in that sense."

By **Kimberly M. S. Cartier** (@AstroKimCartier), Staff Writer

CALL FOR PROPOSALS Scientific Ocean Drilling

The International Ocean Discovery Program (IODP) explores Earth's climate history, structure, mantle/ crust dynamics, natural hazards, and deep biosphere as described at www.iodp.org/science-plan. IODP facilitates international and interdisciplinary research on transformative and societally relevant topics using the ocean drilling, coring, and down-hole measurement facilities JOIDES Resolution (JR), Chikyu, and

Mission-Specific Platforms (MSP). Proposals are being actively sought for all three facilities.

The JR is currently scheduled into the beginning of 2022 (iodp.tamu.edu/scienceops). Due to the recent facility renewal, we plan to schedule JR expeditions through the end of 2024. The JR is expected to operate in the Equatorial and North Atlantic, Gulf of Mexico, Mediterranean, Caribbean, and the Arctic in 2021 and 2022, and to complete its circumnavigation with a return to the eastern Pacific region by 2023, the western Pacific



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in 2023-2024, and potentially the Indian Ocean by the end of 2024. **Proposals for these future operational areas are now needed**.

MSP expeditions are planned to operate once every other year to recover core from targets that are inaccessible by the other facilities (e.g., shallow water, enclosed seas, ice-covered seas). MSP proposals for any ocean are welcomed.

Completely new Chikyu riser proposals (other than CPPs) will not be accepted until publication of a new post-2023 science plan.

We also invite proposals that involve drilling on land and at sea through coordination with the International Continental Drilling Program (ICDP). Investigators are reminded that the interval from first submission to expedition scheduling is on the order of 4-5 years due to the review process and lead time required for scheduling, and that adequate site characterization/site survey data are critical for success. Submission information can be found at www.iodp.org/submitting-proposals.



Submission Deadline: October 1, 2019 • More information: www.iodp.org • Contact: science@iodp.org